Sistemi sa disuretnim energiama

MKA

i klasični i Kvantni Sistemi

MKA

S(E)= klus*(E) ~ klu (* CE)

Ly normaln sistemi U Statisticho - TD

T= (H)NIV P= (DSCE) Sunslin

\$KA

 $Z = \sum_{n=1}^{\infty} e^{-\beta E_n}$ $Z = \sum_{n=1}^{\infty} g(E_n) e^{-\beta E_n}$ -BEn C

F= - KTluZ

 $S = -\left(\frac{\partial F}{\partial T}\right)V_{,N}$ P=- (OF) TIN

Keloricies jena

OYETKUR 242 BY EASTOCT

$$\left(\omega_n = \frac{-s E_n}{2} \right)$$

VKA

$$S = -\left(\frac{2L}{3L}\right)^{1/M}, \quad b = -\left(\frac{2M}{3L}\right)^{1/M}, \quad \langle M \rangle = -\left(\frac{2M}{3L}\right)^{1/M}$$

Razment rabi Erstem oel N wanton to harmonijsnih oscilatora. Pretfustavljapné i da oscilatori Zanemartivo sluso interaguju i da su identični, odrediti Z, F, S i (H) Mogue't vrednosti y vvants omehanië verj osa Latora $\mathcal{E}_{n} = (n + \frac{1}{2}) \mathcal{L}_{\omega}$; n = 0, 1, 2, ...Oscilabori metjurobro NE interagiji

Onalizoran $Z_N = Z_1^N = \left(\sum_{h=0}^{\infty} e^{BE_n}\right)^N$ $Z_1 = \sum_{h=0}^{\infty} e^{BE_n} = \sum_{h=0}^{\infty} e^{BE_n}$ $Z_1 = \sum_{h=0}^{\infty} e^{BE_n} = \sum_{h=0}^{\infty} e^{BE_n}$ $Z_2 = \sum_{h=0}^{\infty} e^{BE_n}$ $= e^{-skw} \sum_{n=0}^{\infty} e^{-snkw}$ $= e^{-sk\omega} \sum_{n=0}^{\infty} (e^{-sk\omega})^n$ 3hw>0 PTw

$$S f w = \frac{f w}{KT} > 0 \implies e^{-skw} < 1$$

$$\sum_{n=0}^{\infty} x^n = \frac{1}{1-x} \quad |x| < 1$$

$$Z_1 = e^{-\frac{skw}{2}} = \frac{1}{1-e^{-skw}}$$

$$= \frac{e^{\frac{skw}{2}}}{e^{-skw}-1} = \frac{1}{2sh(\frac{skw}{2})}$$

$$Z_N = \left[2sh(\frac{skw}{2})\right]^{-N}$$

$$= \sum_{n=0}^{\infty} 2sh(\frac{skw}{2})$$

$$= \frac{1}{2sh(\frac{skw}{2})}$$

$$= \sum_{n=0}^{\infty} 2sh(\frac{skw}{2})$$

(72) = - 1 27 N bj. (72) = - 3lu Z.

Domaci

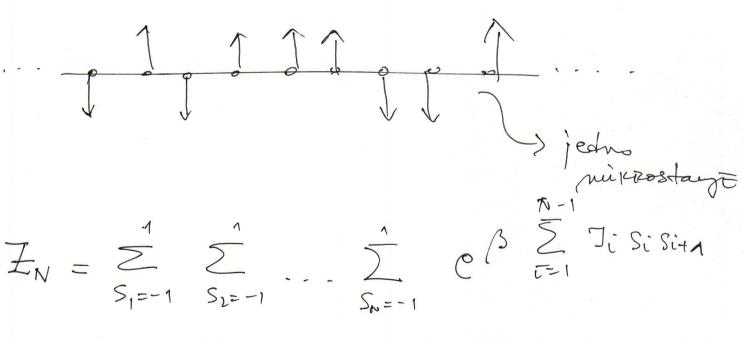
Domaci

Upopediti sa klasserom slucigem (Eada-tak br. 6 i sa resultatom v odeljen "Mikrovanonsni ansamsL" Da li je ovo realon sistem na niskim temperaturama? Ispitati lan S = ? Stedya empegna jednog oserlatopa 2 2 to the 2 Danle, (Je) = N (E)

Jeingov medel. Statisticni sistem je opisas pomoću Hawiltonijana > Delahnje u 22 $\mathcal{H} = -\sum_{i=1}^{n} J_i S_i S_{i+1}$ (*)

got su Ji Konstante, a velicine si mogu imati samo due upednosti! +1 i -1. Odredeti statistione summe, stoSodnu energy i entrov sluiajn nada intens i numerise evopoue NEIVE 1D KRISTALNE PESETUE

Lokalitovan sistem costica



Za jednu oesticu $Z_1 = 2$

Nalazen, E Slabistike Sumet from de u rekurzije jer posto su u pritanje interagnice Testica Nece modi premo obrasca Zn= Zn

$$\frac{7}{N+1} = \frac{7}{S_{1}^{2}-1} \cdot \frac{7}{S_{2}^{2}-1} \cdot \frac{7}{S_{1}^{2}-1} \cdot \frac{7}{S_{1}^$$

3 Nivoc energne oscilatora freuvencye v su oblina $\mathcal{E}_{n} = \left(n + \frac{1}{2} \right) \mathcal{L}_{V}$ (n=0,1,2,.--) Za sistem noji st Sastoji od Noscilatora, Zanlewar hive medpusolent interaneije nadi entromju sistema, nao i vetu izwedju temperature sistema Ti evergijo E. Sistem tretirati Wanto. Totalna energo a

Ez Zii = Z (hit 1) h? $= \sum_{i=1}^{N} n_i h v + \frac{1}{2} N h v$ = MRN + 1 NLV M predstavha broj kvanata hv koje sistem poseduje. Ocigleolno, MZN Primer!

M predstavha broj kvanaba hv kejt sistem posednyt". Ocigleolno, M>N. Primer.

Sta bi Znadibo N = 18 N = 18 N = 19 N = 19 N = 19 N = 19

Kano SE mote Pasporediti ovih M kvamer hi na N oscilatora? Oscilatori su Neradicivi (sistem tretiramo vvantoro), pa poredak nije vazar =) vombinacije sa ponavljanjem

$$[x(E) = (N+M-1)]$$

Za Domodi: M,N>>1, Stirlingova apronsimacoja

$$\frac{1}{T} = \left(\frac{\partial S}{\partial E}\right)_{N_1 V}$$

$$E = \frac{1}{2}NhV + MhV$$

$$M = \frac{E - \frac{1}{2}NhV}{LV} = \frac{E}{hV} - \frac{1}{2}N$$

$$\frac{1}{7} = \frac{K}{hV} eu \left(\frac{N}{M} + 1\right)$$

$$= \frac{K}{hV} tu \left(\frac{N}{E} - \frac{1}{2}N\right)$$

$$\frac{1}{2} = \frac{NhV}{e^{\frac{1}{M}} - 1}$$

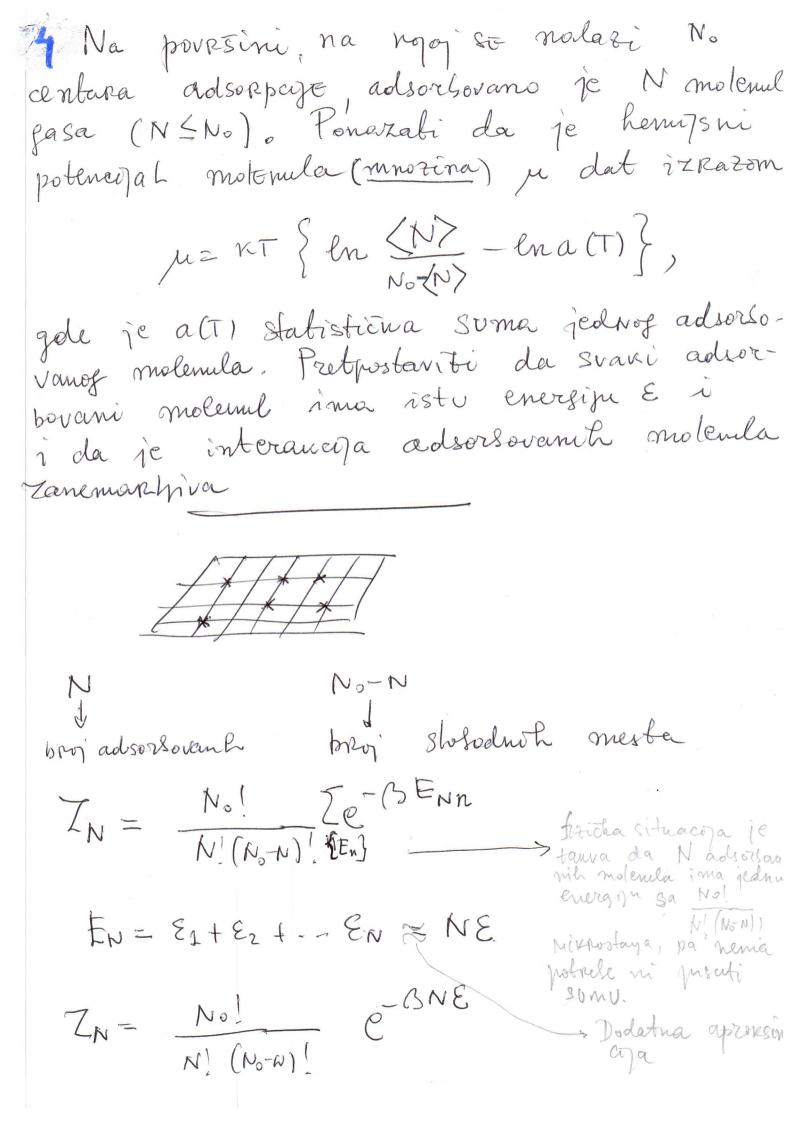
$$= \frac{NhV}{e^{\frac{1}{M}} - 1}$$

$$= \frac{N}{e^{\frac{1}{M}}} - \frac{N}{e^{\frac{1}{M}}} - \frac{N}{e^{\frac{1}{M}}}$$

$$= \frac{N}{e^{\frac{1}{M}}} - \frac{N}{e^{\frac{1}{M}}}$$

$$= \frac{h^{\frac{1}{N}}}{v_{1}} < -1$$

$$= \frac{h^{\frac{1}{N}}}{v_$$



$$Z_{N} = \frac{N_{o}!}{N! (N_{o}N)!} (e^{BE})^{N}$$

$$Q(T) = e^{BE} - \frac{1}{\text{statisticus}} e^{BE}$$

$$Z_{N} = \frac{N_{o}!}{N! (N_{o}-N)!} e^{N}$$

$$Z_{N} = \frac{N_{o}!}{N! (N_{o}-N)!} e^{N}$$

$$Z_{N} = \sum_{N=0}^{N_{o}} \frac{N_{o}!}{N! (N_{o}-N)!} (e^{BE})^{N}$$

$$Z_{N} = \sum_{N=0}^{N_{o}} \frac{N_{o}!}{N! (N_{o}-N)!} (e^{BE})^{N}$$

$$Z_{N} = -kT \ln Z_{N}$$

$$Z_{N} = -kT N_{o} \ln (1+ax)$$

$$\langle N \rangle = -\frac{3N}{3N} = -\frac{3N}{3N} = -\frac{3N}{3N} = -\frac{3N}{3N}$$

$$\langle N \rangle = -\beta \lambda \left(-y \wedge N_0 \right) \frac{1}{1 + a \lambda} \cdot a$$

$$= N_0 \lambda a$$

$$\langle N \rangle = \frac{N_0 \lambda \alpha}{1 + \lambda \alpha}$$

$$\frac{\langle N \rangle}{N_0} = \frac{\lambda \alpha + 1 - 1}{1 + \lambda \alpha} = 1 - \frac{1}{1 + \lambda \alpha}$$

$$\frac{1}{1 + \lambda \alpha} = \frac{1 - \langle N \rangle}{N_0}$$

$$\frac{1}{1 + \lambda \alpha} = \frac{N_0}{N_0 - \langle N \rangle}$$

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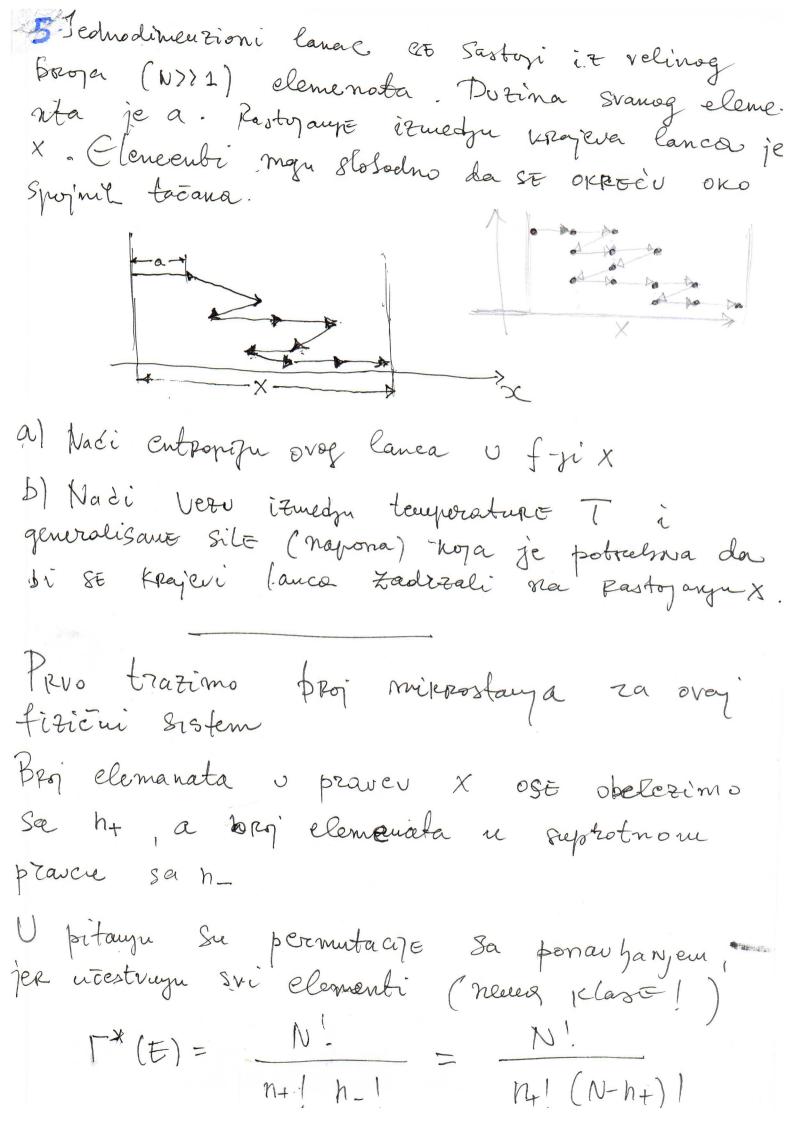
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$$\frac{1}{1$$

$$= \lim_{N \to \infty} \lim_$$



$$n_{+} = \frac{1}{2} \left(N + \frac{x}{\alpha} \right) \qquad n_{-} = \frac{1}{2} \left(N - \frac{x}{\alpha} \right)$$

$$\Gamma^{*}(F) = \frac{1}{2} \left(N + \frac{x}{\alpha} \right) \left[\frac{1}{2} \left(N - \frac{x}{\alpha} \right) \right] \left[\frac{1}{2} \left(N$$

Primetiti it itpara za entropojn. da je 5= S(N,X) = S(E,X) K-Spoyasmi parametar koji se mlaga dx to Spoko H(Zi, Pi, X) -> H(Zi, Pi, X+&X) adjasatski froces kvatistatiehi Micronanous ri dx -> 0 ansambl, dQ=0 U opstem slucaju $Sw = \langle X \rangle dx$ generalisana Adjasat sm proces dE= - < X) dx $dS = \left(\frac{\partial S}{\partial t}\right)_{X} dt + \left(\frac{\partial S}{\partial x}\right)_{n} dx$ ds Zo dx >0

Sistem & Sastoji cel N Neravisn
destica. Svana Testica more da se n
Li Samo na jeotnom od dva energetena
miroa - E, Eo.
a) Naci entropion Si temperatura I
Sistema v f-ji energije E. V proveriti formul
b) Ispitabi u nojoj je oblasti energije
System novemblage i statisticho-terus dinamichou Smishi i v ty oslosti
naci verv itmedyn Eit, van i toploi Kapacitet C.
V
F= N+E0 - N-E0
$=(N+-N-) \in \mathcal{E}_{0}$
= nEo
$V_{+} + N_{-} = N$ $N_{+} - N_{-} = N$
$N_{+} = \frac{n+N}{2}$, $N_{-} = \frac{N-h}{2}$
T*(E) = N!
N+! N-!

$$S = k \left\{ N \ln N - 2N - \frac{n+N}{2} \ln \left(\frac{n+N}{2} \right) - \frac{N-n}{2} \ln \left(\frac{n+N}{2} \right) - \frac{N-n}{2} \ln \left(\frac{N-n}{2} \right) \right\}$$

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retnom popu H. Aus je magnetni moment svans testice je favar da se moze orjentisati samo u dva pravca (pevralelno i antipatralelno popu H), onda su oelgovarque i energrysni vivoi svane certice jet i tjutt. Pretpostavhajuèi da sæ sirtém magnetomh momenata nalazi u toplotnoj ravnoteri sa okulinom, kao i da magnetni momenti medyu-Solons ne interagijn, naci spednju vkednost uniepnog magnetnog momenta hjegova dispiroju, statistichu Sumu sistema i valopochu j mu Verovatuoiea malastrya v stagn (Comt odgova-na energy a En Un = C-15En Wh = Comt odgova-Magnetin monenti imagn zanemar hive nonebrohn energym, pa sistem moremo Smortrati pribliono Conalito vanon

$$\begin{aligned}
\overline{I} &= 2ch \left(\beta_{M}H\right) \\
\overline{I} &= I^{N} = 2^{N}ch^{N} \left(\beta_{M}H\right) \\
\overline{I} &= I^{N} = 2^{N}ch^{N} \left(\beta_{M}H\right) \\
F &= NI lu \overline{I} = ... = -NKT lu \left(2ch \left(\beta_{M}H\right)\right) \\
S &= -\left(\frac{3F}{2T}\right)_{V,N} = ... = NK \left\{lu \left(2ch \left(\beta_{M}H\right)\right) - \beta_{M}H \left(lh \left(\beta_{M}H\right)\right)\right\} \\
Magnethi moment sistema

$$\overline{M} &= \sum_{i=1}^{N} \overline{I_{i}} \\
\overline{M} &= \sum_{i=1}^{N} \overline{I_{i}} \\
\overline{I_{i}} &= \sum_{i=1}^$$$$

(M) > = Nuth (But) = BZ² / H²? - BZ² (M)². D(M) = 1 D(M) = XM

Magneterna susceptibiliset

D(M) = 1 D(M) = NM [th BMH]

H $= \frac{N_M}{p} \frac{1}{ch^2(p_M)} p_M$ $D(M') = N\mu^2 \frac{1}{Ch^2(\beta MH)}$ Calopicha I-na storger U= F+TS = ... = - (M), # La alternativno peravage, videti Hirku luho et al, str. 93,

Spentar vrednosti 8 Sistem ima nedegenerisani energija Ex=KE (K=0,1,2,...). Odreduti sredyn energym sistema. Raditi u formalit mu vanon-Skog ansamsLa.

Zadaho je
$$E_k = KE$$

$$\mathcal{Z} = \sum_{K=0}^{\infty} e^{-15KE} = \sum_{K=0}^{\infty} (e^{-15E})^K = \frac{1}{1-e^{5E}}$$

$$(E) = -\frac{\partial}{\partial s} \ln \frac{1}{1-\bar{e}^s \epsilon} = \frac{\partial}{\partial s} \ln (1-\bar{e}^s \epsilon)$$

$$\frac{1}{1-\tilde{\epsilon}^{56}} = \frac{\epsilon}{1-\tilde{\epsilon}^{56}-1}$$

Domacil: Naci Cy La ovoj fitich sistem

$$C_{Y} = \frac{\partial \langle E \rangle}{\partial T} = \frac{\partial \langle E \rangle}{\partial B} = \frac{1}{RT} \frac{\partial \langle E \rangle}{\partial B}$$

$$= kB^2 \frac{e^2}{(e^{BE}-1)^2}$$

Iomaci 2: Naci izpat da entroppe ovog fitiolog sistema.

$$S = -(3f)_{y} = -3f = -3f = -3f = -12f$$

$$\frac{\partial F}{\partial S} = \frac{\epsilon}{B} \frac{1}{e^{B} \epsilon^{-1}} + \frac{1}{B^{2}} lu \left(\frac{1}{1 - \epsilon^{-B} \epsilon}\right)$$

$$= \frac{\epsilon}{B} \frac{1}{e^{B} \epsilon^{-1}} + \frac{1}{B^{2}} lu \left(1 - \epsilon^{-B} \epsilon\right)$$

D(E)= 92 to C + 9, 9, 62 c -16/6/6) +8,9,6,0 (4/6) + 9260 C -276, C +15,906,620 (4/6) - 3266 D(4) = (9,00 +3.00)(9,5,00 +52520 +52520) - 9,5,00 +23,25,5,6 -36,44) +9,2200 -2052 (E2)= Siere seitsepe Cre Z= 9, ent, + 92 cot (E)= 9, E, e'SE, + g. e.c. SEr F2/3Ez (-9, 2-2/2(E1-E2) 99 9 PS (E1-E2) -3,9,0 >(6,162) (Ez +6,2 - 2E, Ez) 9,92 a c s(E,+E2) 31 6 2/2 + 22/2 6 2(E/ter) + 22 C -2/5 = 2 Naci flustración enceptre Za sistem se somo 9,92 a2 C-Ba (9,+9,002)2 W ビューでユー b EztEz=a 92 E-20E1+29192 C-3(E1+E2) + 92 C-15E2 (E1-E2=0 3, 9, (+1-62)2 E-2 (E)+62) C2 = R-5

144

Principles of tquilibrium Statistical

3.7.2 Canonical Ensemble

The partition function for ideal Bose gas and ideal Fermi gas can be written as

$$Z = \sum_{\{n_i\}} \exp[-\beta(n_1\epsilon_1 + n_2\epsilon_2 + ...)]$$
 (3.135)

So, the average number of particles in a state r is

$$\langle n_r \rangle = \frac{\sum_{n_1, n_2, \dots} n_r e^{-\beta(n_1 \epsilon_1 + n_2 \epsilon_2 + \dots + n_r \epsilon_r + \dots}}{\sum_{n_1, n_2, \dots} e^{-\beta(n_1 \epsilon_1 + n_2 \epsilon_2 + \dots + n_r \epsilon_r + \dots}}$$
 (3.136)

or,

$$\langle n_r \rangle = -\{1/(\beta Z)\}(\partial Z/\partial \epsilon_r) = -(1/\beta)[\partial(\ln Z)/\partial \epsilon_r]$$
 (3.137)

The formula (3.136) can be used to derive the mean occupation of the levels in a quantum gas. But, in order to emphasize the crucial differences between the Fermi gas, Bose gas (and photon gas) we shall calculate $< n_r >$ directly from its fundamental definition (3.135) which can be reexpressed as

$$< n_r> = \frac{[\{\sum_{n_r} n_r e^{-\beta n_r \epsilon_r}\} \{\sum_{n_1,n_2,\dots}^{(r)} e^{-\beta (n_1 \epsilon_1 + n_2 \epsilon_2 + \dots + n_r \epsilon_r + \dots)}\}]}{[\{\sum_{n_r} e^{-\beta n_r \epsilon_r}\} \{\sum_{n_1,n_2,\dots}^{(r)} e^{-\beta (n_1 \epsilon_1 + n_2 \epsilon_2 + \dots + n_r \epsilon_r + \dots)}\}]}$$

where the symbol $\sum_{n_1,n_2,...}^{(r)}$ denotes sum over all states except the r-th. Note that, the allowed values of n_r depend on the nature of the particles, i.e., whether these are Fermions or Bosons. Moreover, the upper limits of the n's in the summations are, in general, correlated because of the condition $\sum_i n_i = N$.

However, the formula (3.135) gets simplified in the case of the photon gas where each level can accommodate any number of particles and total number of particles need not be conserved. Then, each of the sums in (3.135) extends from 0 to ∞ , independent of each other. Consequently, the second factors within the brackets in the numerator and denominator cancel each other and we are left with the simpler expression

$$\langle n_r \rangle = \left[\sum_{n_r} n_r e^{-\beta n_r \epsilon_r} \right] / \left[\sum_{n_r} e^{-\beta n_r \epsilon_r} \right]$$

$$= -(1/\beta)(\partial/\partial \epsilon_r) \left[\ln \left\{ \sum_{n_r=0}^{\infty} \left(e^{-\beta \epsilon_r} \right)^{n_r} \right\} \right] = (1/\beta)(\partial/\partial \epsilon_r) [\ln(1 - e^{-\beta \epsilon_r})]$$

i.e.,

$$\langle n_r \rangle = [e^{\beta \epsilon_r} - 1]^{-1}$$
 (3.138)

which is the Planck distribution for the photon gas.

As we shall demonstrate now, one has to proceed more carefully in the case of ideal quantum gases where the total number of particles is conserved. The symbol

$$Z_r(N) = \sum_{n_1, n_2, \dots}^{(r)} e^{-\beta(n_1\epsilon_1 + n_2\epsilon_2 + \dots)}$$
 (3.139)

denotes the partition function of a system of N particles which are distributed over all the states except the r-th, i.e., N particles are accounted for before the remaining sum over the states of the r-th cells are carried out.

•Ideal Fermi Gas

Since for Fermions, 0 and 1 are the only two allowed values of n_r

$$< n_r > = \frac{0 + e^{-\beta \epsilon_r} Z_r(N-1)}{Z_r(N) + e^{-\beta \epsilon_r} Z_r(N-1)} = \left[\frac{Z_r(N)}{Z_r(N-1)} e^{\beta \epsilon_r} + 1 \right]^{-1}$$

But, for $\Delta N \ll N$,

$$\ln Z_r(N - \Delta N) \simeq \ln Z_r(N) - \alpha \Delta N \tag{3.140}$$

where $\alpha = \partial(\ln Z_r)/\partial N = -\mu/(k_B T)$. Taking $\Delta N = 1$, we have

$$Z_r(N-1) \simeq e^{-\alpha} Z_r(N) \tag{3.141}$$

and, hence,

$$\langle n_r \rangle = \left[e^{\alpha + \beta \epsilon_r} + 1 \right]^{-1} \tag{3.142}$$

which is identical to the equation (3.123).

In the above derivation we have made two assumptions, (i) that $\alpha_r = \partial(\ln Z_r)/\partial N = \alpha$, independent of r; and (ii) one can truncate the series in (3.139) after the second term even when $\Delta N = 1$. We can now derive conditions required to be satisfied for the validity of these two assumptions. Since,

$$Z(N) = Z_r(N)[1 + e^{-\alpha - \beta \epsilon_r}]$$

taking derivatives of the logarithms of both sides with respect to N

$$\alpha = \alpha_r - \langle n_r \rangle (\partial \alpha / \partial N)$$

Therefore, in order that $\alpha = \alpha_r$, we must have $(\partial \alpha/\partial N) < n_r > \ll \alpha$ and since $< n_r > \leq 1$, the required condition to be satisfied is $(\partial \alpha/\partial N) \ll \alpha$, i.e., if the chemical potential does not change significantly on additing a particle. This condition is normally satisfied by a macroscopically large system of ideal Fermi gas. The justification for the assumption (ii) can be given along the same line as some of the arguments presented in sections 3.3 and 3.4 in the context of classical systems.

•Ideal Bose Gas

Since, for Bosons, all the integers 0, 1, 2, 3, ... are allowed values of n_r , with the abbreviation $y = \exp(-\beta \epsilon_r)$,

$$< n_r > = \frac{0 + y Z_r(N-1) + 2y^2 Z_r(N-2) + 3y^3 Z_r(N-3) + \dots}{Z_r(N) + y Z_r(N-1) + y^2 Z_r(N-2) + y^3 Z_r(N-3) + \dots}$$

$$= \frac{e^{-(\alpha + \beta \epsilon_r)} + 2e^{-2(\alpha + \beta \epsilon_r)} + 3e^{-3(\alpha + \beta \epsilon_r)} + \dots}{1 + e^{-(\alpha + \beta \epsilon_r)} + e^{-2(\alpha + \beta \epsilon_r)} + e^{-3(\alpha + \beta \epsilon_r)} + \dots},$$

or,

$$\langle n_r \rangle = [e^{\alpha + \beta \epsilon_r} - 1]^{-1} \tag{3.143}$$

•Ideal Boltzmann Gas

For a given set $\{n_1, n_2, ...\}$, there are $N!/(n_1!n_2!...)$ possible ways in which the N particles can be put into the given single-particle states. So, in contrast to the equation (3.134) for the partition function for ideal Bose and fermi gases, the partition function for ideal Boltzmann gas is obtained from

$$Z = \sum_{n_1, n_2, \dots} [N!/(n_1!n_2!\dots)]e^{-\beta(n_1\epsilon_1 + n_2\epsilon_2 + \dots)}$$
$$= \sum_{n_1, n_2, \dots} [N!/(n_1!n_2!\dots)](e^{-\beta\epsilon_1})^{n_1}(e^{-\beta\epsilon_2})^{n_2}\dots$$

i.e.,

$$Z = (e^{-\beta\epsilon_1} + e^{-\beta\epsilon_2} + \dots)^N$$

Therefore, $\ln Z = N \ln(\sum_r e^{-\beta \epsilon_r})$ which leads to

$$\langle n_r \rangle = Ne^{-\beta\epsilon_r} / (\sum_r e^{-\beta\epsilon_r})$$
 (3.144)

which is identical to classical Maxwell-Boltzmann distribution.

Solve problem 3.23.

3.7.3 Grand-Canonical Ensemble

Just as in case of canonical ensemble, we can now show that in case of the grand canonical ensemble,

$$\langle n_r \rangle = -(1/\beta)[\partial(\ln Z_G)/\partial \epsilon_r]$$

In this ensemble, the derivation of $\langle n_r \rangle$ is the simplest.

•Ideal Fermi Gas

The grand partition function is given by

$$Z_G = \prod_i \left[\sum_{n_r=0}^1 e^{-eta(\epsilon_r - \mu)n_r} \right]$$

or, ln $Z_G = \sum_r \ln[1 + e^{\beta(\mu - \epsilon_r)}]$ and, hence,

$$\langle n_r \rangle = [e^{\beta(\epsilon_r - \mu)} + 1]^{-1}$$
 (3.145)

•Ideal Bose Gas

In this case

$$Z_G = \prod_i \left[\sum_{n_r=0}^{\infty} e^{-\beta(\epsilon_r - \mu)n_r} \right]$$

i.e., $\ln Z_G = -\sum_r \ln[1 - e^{\beta(\mu - \epsilon_r)}]$ and, hence,

$$\langle n_r \rangle = [e^{\beta(\epsilon_r - \mu)} - 1]^{-1}$$
 (3.146)

Solve problem 3.24 - 3.26.

3.8 Quantum Systems; Density Operator

In quantum mechanics, say, in the Dirac formulation, the completely prepared states are represented by their "ket" vectors $|\lambda\rangle$. But, in quantum statistical mechanics, the preparations are always incomplete and, hence, a system is not described by a well-defined ket. This situation is the quantum counterpart of the incomplete knowledge about the micro-states in classical statistical mechanics, where we do not know in which micro-state the system has been prepared. Naturally, just as we assign a probability P_r with each micro-state r in classical statistical mechanics, we assign a probability

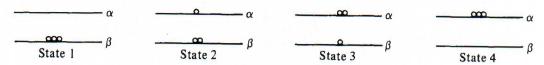


Fig. 4.1. The four states of a three-particle system with two single particle states.

large collection of particles. An example is the small amplitude vibrational modes of a solid. These modes are called *phonons*. Another example is the *occupation numbers* for quantum mechanical systems composed of non-interacting particles.

In this chapter, we will consider phonons, occupation numbers, classical ideal gases, and a number of other examples to illustrate how the factorization method is applied.

Chardler Introduction to modern statistical

4.1 Occupation Numbers

Physics

The first step in analyzing any model involves the classification of microstates. The state of a quantum system can be specified by the wavefunction for that state, $\Psi_{\nu}(r_1, r_2, \ldots, r_N)$. Here, Ψ_{ν} is the ν th eigensolution to Schrödinger's equation for an N-particle system. If the particles are non-interacting (i.e., ideal), then the wavefunction can be expressed as a symmetrized* product of single particle wavefunctions. Let us denote these single particle wavefunctions as $\phi_1(r)$, $\phi_2(r)$, ..., $\phi_i(r)$, For a particular state, say ν , $\Psi_{\nu}(r_1,\ldots,r_N)$ will be a symmetrized product containing n_1 particles with the single particle wavefunction ϕ_1 , n_2 particles with the single particle wavefunction ϕ_2 , and so on. These numbers, $n_1, n_2, \ldots, n_i, \ldots$ are called the occupation numbers of the first, second, ...jth, ... single particle states. If the N particles are indistinguishable—as quantum particles are—then a state, v, is completely specified by the set of occupation $(n_1, n_2, \ldots, n_i, \ldots)$ since any more detail would distinguish between the n_i particles in the jth single particle state.

For example, consider three particles (denoted in Fig. 4.1 by circles) which can exist in one of two single particle states, α and β . All the possible states for this three-particle system are exhibited in Fig. 4.1. In terms of occupation numbers, state 1 has $n_{\alpha} = 0$, $n_{\beta} = 3$; state 2 has $n_{\alpha} = 1$, $n_{\beta} = 2$; and so on. Notice that an occupation number is a collective variable in the sense that its value depends upon the instantaneous state of all the particles.

Let us now express the total number of particles and the total

^{*} For Fermi particles, the product is antisymmetric; for Bose particles the product is symmetric.

energy in terms of the occupation numbers. Let

$$v = (n_1, n_2, ..., n_j, ...) = v$$
th state.

Then

$$N_{\nu} = \sum_{j} n_{j} = \text{total number of particles in the } \nu \text{th state.}$$

Let ε_j be the energy of the jth single particle state. Then,

$$E_{\nu} = \sum_{i} \varepsilon_{i} n_{i} = \text{energy in the } \nu \text{th state.}$$

Particles with half-integer spin obey an exclusion principle*: $n_j = 0$ or 1, only. Such particles are called *fermions* and the statistics associated with $n_j = 0$ or 1 is called *Fermi-Dirac* statistics.

Particles with integer spin obey Bose-Einstein statistics: $n_j = 0, 1, 2, 3, \ldots$ These particles are called bosons.

4.2 Photon Gas

As an example of how we use occupation numbers, consider the photon gas—an electromagnetic field in thermal equilibrium with its container. We want to describe the thermodynamics of this system. From the quantum theory of the electromagnetic field, it is found that the Hamiltonian can be written as a sum of terms, each having the form of a Hamiltonian for a harmonic oscillator of some frequency. The energy of a harmonic oscillator is $n\hbar\omega$ (zero point energy omitted), where $n = 0, 1, 2, \ldots$ Thus, we are led to the concept of photons with energy $\hbar\omega$. A state of the free electromagnetic field is specified by the number n for each of the "oscillators," and n can be thought of as the number of photons in a state with single "particle" energy $\hbar\omega$.

Photons obey Bose-Einstein statistics: $n = 0, 1, 2, \ldots$ The canonical partition function is thus

$$e^{-\beta A} = Q = \sum_{\mathbf{v}} e^{-\beta E_{\mathbf{v}}} = \sum_{\substack{n_1, n_2, \dots, n_j, \dots \\ = 0}}^{\infty} e^{-\beta (n_1 \varepsilon_1 + n_2 \varepsilon_2 + \dots + n_j \varepsilon_j + \dots)},$$

where we have used the occupation number representation of E_{ν} , and denoted $\hbar\omega_i$ by ε_i . Since the exponential factors into independ-

^{*} The requirement that the N-particle wavefunction be an antisymmetric product implies the exclusion principle.

ent portions, we have

$$Q = \prod_{j} \left[\sum_{n_{j}=0}^{\infty} e^{-\beta n_{j} \varepsilon_{j}} \right].$$

The term in brackets is just a geometric series, thus

$$Q(\text{photon gas}) = \prod_{j} \left[\frac{1}{1 - e^{-\beta \varepsilon_j}} \right].$$

From this formula, we can obtain all the properties we want since $Q = e^{-\beta A}$. One quantity that is particularly interesting is the average value of the occupation number of the jth state, $\langle n_j \rangle$. In the canonical ensemble

$$\langle n_{j} \rangle = \frac{\sum_{v} n_{j} e^{-\beta E_{v}}}{\sum_{v} e^{-\beta E_{v}}} = \frac{\sum_{n_{1}, n_{2}, \dots} n_{j} e^{-\beta (n_{1} \varepsilon_{1} + \dots + n_{j} \varepsilon_{j} + \dots)}}{Q}$$

$$= \left[\frac{\partial}{\partial (-\beta \varepsilon_{j})} \sum_{n_{1}, n_{2}, \dots} e^{-\beta (n_{1} \varepsilon_{1} + \dots + n_{j} \varepsilon_{j} + \dots)} \right] / Q$$

$$= \frac{\partial \ln Q}{\partial (-\beta \varepsilon_{j})}.$$

Returning to our formula for Q we thus have

$$\langle n_j \rangle = + \frac{\partial}{\partial (-\beta \varepsilon_j)} \left\{ \sum_j - \ln (1 - e^{-\beta \varepsilon_j}) \right\}$$

= $e^{-\beta \varepsilon_j} / [1 - e^{-\beta \varepsilon_j}]$

or

$$\langle n_j \rangle = [e^{\beta \varepsilon_j} - 1]^{-1},$$

which is called the Planck distribution.

Exercise 4.1 For the photon gas derive a formula for the correlation function $\langle \delta n_i \delta n_j \rangle$ where $\delta n_i = n_i - \langle n_i \rangle$.

Exercise 4.2* Use the formula for $\langle n_j \rangle$ to show that the energy density of a photon gas is σT^4 , where σ is a constant, $(\pi^2 k_B^4/15\hbar^3 c^3)$. [Hint: You will need to construct a formula for the number of standing wave solutions to use the wave equation for waves in a three-dimensional cavity and with frequency between ω and $\omega + d\omega$.]